

Hyperbolic case : similarly

$$\sum_n \ddot{a}_n(t) v_n(x) + \sum_n \lambda_n a_n(t) v_n(x) = \sum_n b_n(t) v_n(x)$$

so by orthogonality

$$\ddot{a}_n(t) + \lambda_n a_n(t) = b_n(t)$$

This time we use the Laplace transform method:

$$s^2 \hat{a}_n - s a_n(0) - a_n'(0) + \lambda_n \hat{a}_n = \hat{b}_n(s)$$

$$\Rightarrow \hat{a}_n(s) = \frac{\hat{b}_n(s) + s a_n(0) + a_n'(0)}{s^2 + \lambda_n}$$

Now, the λ_n are positive \Rightarrow the inverse Laplace transform (see Table) is

$$a_n(t) = \frac{1}{\sqrt{\lambda_n}} \int_0^t b_n(t') \sin(\sqrt{\lambda_n}(t-t')) dt' \\ + a_n(0) \cos(\sqrt{\lambda_n} t) + a_n'(0) \frac{1}{\sqrt{\lambda_n}} \sin(\sqrt{\lambda_n} t)$$

\Rightarrow The general solution of the problem becomes

$$u(x,t) = \sum_{n=0}^{\infty} v_n(x) \cdot \left[\frac{1}{\sqrt{\lambda_n}} \int_0^t b_n(t') \sin(\sqrt{\lambda_n}(t-t')) dt' \right. \\ \left. + a_n(0) \cos(\sqrt{\lambda_n} t) + a_n'(0) \frac{1}{\sqrt{\lambda_n}} \sin(\sqrt{\lambda_n} t) \right]$$

$$\text{but with } b_n(t') = \int_a^b \frac{F(x',t') v_n(x') r(x') dx'}{\|v_n\|^2}$$

we get

$$u(x,t) = \sum_n \left[a_n(0) \cos(\sqrt{\lambda_n} t) + a_n'(0) \frac{1}{\sqrt{\lambda_n}} \sin(\sqrt{\lambda_n} t) \right] \cdot v_n(x) \\ + \int_0^t \int_a^b F(x',t') G(x, x'; t, t') dx' dt'$$

with

$$G(x, x'; t, t') = \sum_{n=0}^{\infty} \frac{1}{\sqrt{A_n}} \sin(\sqrt{A_n}(t-t')) \frac{v_n(x)v_n(x')r(x')}{\|v_n\|^2}$$

↳ the wave kernel

Example of the bridge

Recall: $u_{tt} - c^2 u_{xx} = \sin\left(\frac{2\pi x}{L}\right) \cos(\omega t)$

$u = 0$ at both ends

$$u_t(x, 0) = u(x, 0) = 0$$

⇒ Eigenmodes/values of spatial homogeneous pb:

$$\begin{cases} v_n(x) = \sin\left(\frac{n\pi x}{L}\right) \\ A_n = \frac{n^2 \pi^2 c^2}{L^2} \end{cases}$$

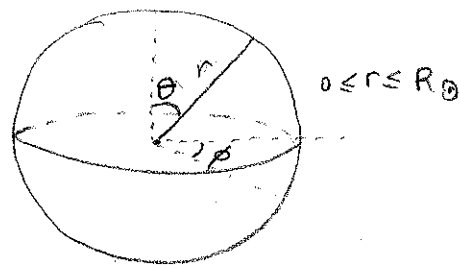
then
$$u(x, t) = \sum \left[a_n(0) \cos\left(\frac{n\pi c t}{L}\right) + a_n'(0) \frac{L}{n\pi c} \sin\left(\frac{n\pi c t}{L}\right) \right] \sin\left(\frac{n\pi x}{L}\right) + \int_0^t \int_0^L F(x', t') G(x, x'; t, t') dx' dt'$$

Fitting this to ICs ⇒ $a_n(0) = a_n'(0) = 0$

$$\begin{aligned} u(x, t) &= \int_0^t \int_0^L \sin\left(\frac{2\pi x'}{L}\right) \cos(\omega t') \sum_{n=0}^{\infty} \frac{L}{n\pi c} \sin\left(\frac{n\pi c}{L}(t-t')\right) \frac{\sin\left(\frac{n\pi x}{L}\right) \sin\left(\frac{n\pi x'}{L}\right)}{\frac{L}{2}} \\ &= \int_0^t dt' \cos(\omega t') \frac{L}{2\pi c} \sin\left(\frac{2\pi c}{L}(t-t')\right) \sin\left(\frac{2\pi x}{L}\right) \\ &= \int_0^t \frac{dt'}{2} \left[\sin\left(\omega t' + \frac{2\pi c}{L}(t-t')\right) - \sin\left(\omega t' - \frac{2\pi c}{L}(t-t')\right) \right] \sin\left(\frac{2\pi x}{L}\right) \frac{L}{2\pi c} \\ &= \frac{1/2}{\omega + \frac{2\pi c}{L}} \left[\cos\left(\frac{2\pi c t}{L}\right) - \cos \omega t \right] - \frac{1/2}{\omega - \frac{2\pi c}{L}} \left[\cos\left(\frac{2\pi c}{L}\right) - \cos \omega t \right] \frac{L}{2\pi c} \\ &= \frac{1}{\omega^2 - \frac{4\pi^2 c^2}{L^2}} \left[\cos\left(\frac{2\pi c t}{L}\right) - \cos \omega t \right] \sin\left(\frac{2\pi x}{L}\right) \checkmark \cdot \sin\left(\frac{2\pi x}{L}\right) \end{aligned}$$

Example of application to 3D problems

Solar oscillations



- The Sun is a spherical ball of self-gravitating gas.
- The outer layer is convecting; the convective motions excite sound waves which propagate throughout the interior. The sound waves satisfy the approximate equation

$$\frac{\partial^2 p}{\partial t^2} = c_s^2(r) \nabla^2 p$$

p = pressure

c_s = sound speed, assumed to depend only on radius r

∇^2 = Laplacian in spherical coordinates r, θ, ϕ

$$\Rightarrow \frac{\partial^2 p}{\partial t^2} = c_s^2(r) \left[\frac{1}{r^2} \frac{\partial}{\partial r} \left(r^2 \frac{\partial p}{\partial r} \right) + \frac{1}{r^2 \sin \theta} \frac{\partial}{\partial \theta} \left(\sin \theta \frac{\partial p}{\partial \theta} \right) + \frac{1}{r^2 \sin^2 \theta} \frac{\partial^2 p}{\partial \phi^2} \right]$$

We will assume that the boundary conditions are

$$\begin{cases} p(0, \theta, \phi, t) < +\infty \\ p(R_0, \theta, \phi, t) = 0 \end{cases}$$

Eigenmodes

Separation of variables \Rightarrow assume

$$p(r, \theta, \phi, t) = A(r) B(\theta) C(\phi) D(t)$$

$$\Rightarrow \underbrace{\frac{1}{D} \frac{d^2 D}{dt^2}}_{\text{a function of time only}} = c_s^2(r) \underbrace{\left[\frac{1}{A} \frac{1}{r^2} \frac{d}{dr} \left(r^2 \frac{dA}{dr} \right) + \frac{1}{B} \frac{1}{r^2 \sin \theta} \frac{d}{d\theta} \left(\sin \theta \frac{dB}{d\theta} \right) + \frac{1}{C} \frac{1}{r^2 \sin^2 \theta} \frac{d^2 C}{d\phi^2} \right]}_{\text{a function of space only}}$$

$$= -\omega^2$$

\uparrow expect oscillations so select a negative constant

$$\Rightarrow \left\{ \begin{array}{l} \frac{d^2 D}{dt^2} = -\omega^2 D \end{array} \right. \quad (1)$$

$$\left\{ \begin{array}{l} \frac{1}{A} \frac{d}{dr} \left(r^2 \frac{dA}{dr} \right) + \frac{1}{B \sin \theta} \frac{d}{d\theta} \left(\sin \theta \frac{dB}{d\theta} \right) + \frac{1}{C \sin^2 \theta} \frac{d^2 C}{d\phi^2} = -\frac{\omega^2 r^2}{c^2(r)} \end{array} \right. \quad (2)$$

take (2)

$$\Rightarrow \underbrace{\frac{1}{A} \frac{d}{dr} \left(r^2 \frac{dA}{dr} \right) + \frac{\omega^2 r^2}{c^2(r)}}_{\text{A function of } r \text{ only}} = - \underbrace{\left[\frac{1}{B \sin \theta} \frac{d}{d\theta} \left(\sin \theta \frac{dB}{d\theta} \right) + \frac{1}{C \sin^2 \theta} \frac{d^2 C}{d\phi^2} \right]}_{\text{A function of } \theta \text{ and } \phi \text{ only}}$$

$$= -\lambda$$

↑ another constant

$$\text{so } \left\{ \begin{array}{l} \frac{d}{dr} \left(r^2 \frac{dA}{dr} \right) + \lambda A = -\frac{\omega^2 r^2}{c^2(r)} A \quad (3) \\ \frac{1}{B \sin \theta} \frac{d}{d\theta} \left(\sin \theta \frac{dB}{d\theta} \right) + \frac{1}{C \sin^2 \theta} \frac{d^2 C}{d\phi^2} = \lambda \quad (4) \end{array} \right.$$

take (4), multiply by $\sin^2 \theta \rightarrow$

$$\underbrace{-\lambda \sin^2 \theta + \frac{1}{B} \sin \theta \frac{d}{d\theta} \left(\sin \theta \frac{dB}{d\theta} \right)}_{\text{A function of } \theta \text{ only}} = - \underbrace{\frac{1}{C} \frac{d^2 C}{d\phi^2}}_{\text{A function of } \phi \text{ only}}$$

$$= m^2$$

↑ a constant.

Since we expect 2π -periodic functions in ϕ , we can straightaway recognize that constant to be m^2 , with m integer.

So $\left\{ \begin{array}{l} \frac{d^2 C}{d\phi^2} = -m^2 C \quad (5) \end{array} \right.$

$\left\{ \begin{array}{l} \sin\theta \frac{d}{d\theta} \left(\sin\theta \frac{dB}{d\theta} \right) - m^2 B = \lambda \sin^2\theta B \quad (6) \end{array} \right.$

\Rightarrow (1), (3), (5) and (6) are the 4 equations representing the eigenmodes in the t, r, θ and ϕ variables.

Solution in ϕ : naturally $C_m(\phi) = \left\{ \begin{array}{l} \cos m\phi \\ \sin m\phi \end{array} \right\}$

Solution in θ : The θ -mode depends on the value of m

Trick Let $\cos\theta = \mu$
then

$$\frac{d}{d\theta} = \frac{d}{d\mu} \frac{d\mu}{d\theta} = -\sin\theta \frac{d}{d\mu}$$

so (6) becomes

$$(1-\mu^2) \frac{d}{d\mu} \left((1-\mu^2) \frac{dB}{d\mu} \right) - m^2 B = \lambda (1-\mu^2) B$$

$$\rightarrow \frac{d}{d\mu} \left((1-\mu^2) \frac{dB}{d\mu} \right) - \frac{m^2}{1-\mu^2} B = \lambda B$$

The solutions to this equation with $\lambda = -l(l+1)$ are the Legendre functions

$$\left\{ P_l^m(\mu) \right\}_{l, m \text{ integers}}$$

So $B_l^m(\theta) = \left\{ P_l^m(\cos\theta) \right\}_{m, l \text{ integers}}$

Some properties of Legendre functions:

- $P_l^m(x) = (-1)^{|m|} (1-x^2)^{\frac{|m|}{2}} \frac{d^{|m|} P_l(x)}{dx^{|m|}}$

where $P_l(x)$ is the Legendre Polynomial of order l

- $P_l(x) = P_l^0(x)$

- $P_0(x) = 1$

$$P_1(x) = x$$

$$(n+1) P_{n+1}'(x) = (2n+1)x P_n'(x) - n P_{n-1}'(x)$$

- So $P_l^m(x) = 0 \quad \forall |m| > l$

Note: The combination of the two angular functions are usually called spherical harmonics and noted $Y_l^m(\theta, \phi)$

e.g. $Y_2^1(\theta, \phi) \propto P_2^1(\cos\theta) \cos\phi$

although standard conventions usually use

$$Y_l^m(\theta, \phi) \propto P_l^m(\cos\theta) e^{i\phi}$$

$$Y_l^m(\theta, \phi) \propto P_l^m(\cos\theta) e^{im\phi}$$

Solution in r

$$\frac{d}{dr} \left(r^2 \frac{dA}{dr} \right) - l(l+1)A = -\frac{\omega^2 r^2}{c_s^2(r)} A$$

→ clearly, this equation cannot be solved directly without knowledge of $c_s^2(r)$. However, we can already say a lot about the solution by inspection.

+ this is a Sturm-Liouville problem with

$$p(r) = r^2$$

$$q(r) = -l(l+1)$$

$$w(r) = \frac{r^2}{c_s^2(r)} \quad \leftarrow \text{weight function}$$

$$A = \omega^2$$

Let's take the simplest possible example, that of constant sound-speed $c_s(r)$. Then

$$\frac{d}{dr} \left(r^2 \frac{dA}{dr} \right) - l(l+1)A = - \frac{\omega^2}{c_s^2} r^2 A$$

This is actually the equation for a spherical Bessel function.

Indeed, spherical Bessel functions satisfy

$$\frac{d}{dx} \left(x^2 \frac{dA}{dx} \right) - l(l+1)A + x^2 A = 0. \Rightarrow \text{as above if}$$

$$x = \frac{\omega}{c} r$$

$$A_l(x) = \begin{cases} j_l(x) \\ y_l(x) \end{cases}$$

The $y_l(x)$ functions are singular at $x=0 \rightarrow$ discard.

$$\text{At the surface, } A_l(x) = 0 \Rightarrow A_l\left(\frac{\omega}{c} R_*\right) = 0.$$

\Rightarrow this means that $\frac{\omega}{c} R_*$ are zeros of the Bessel $j_l(x)$ function. There is an infinite number of them, noted z_n (the n -th zero of the $j_l(x)$ function)

$$\Rightarrow \omega_{nl} = \frac{z_n}{R_*} c_s.$$

\Rightarrow Finally, putting everything together we have

$$p(r, \theta, \varphi, t) = \sum_{n, l, m} A_{nl}(r) B_l^m(\theta) C_m(\varphi) D_{nlm}(t) \text{ where}$$

$$A_{nl}(r) = j_l\left(z_n \frac{r}{R_*}\right)$$

$$C_m(\varphi) B_l^m(\theta) = Y_l^m(\theta, \varphi) \leftarrow \text{spherical harmonic}$$

$$D_{nlm}(t) = a_{nlm} \cos(\omega_{nl} t) + b_{nlm} \sin(\omega_{nl} t)$$

$$\text{where } \omega_{nl} = \frac{z_n}{R_*} c_s$$